Developing and Testing an Improved Description of Diffusion For Use in Modeling Planetary Atmospheres

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The objective of this experiment was to create an improved conduction code for the Global Ionoshpere-Thermosphere Model (GITM) used by NASA. GITM used an implicit solver to the diffusion equation, which is an unconditionally stable scheme, but it sacrifices a lot of accuracy, being only first order accurate. The proposed replacement scheme is the Crank-Nicolson scheme, which is a hybridization of an implicit and explicit solver and has second order accuracy. This hybrid scheme combines the accuracy of the explicit solver, while theoretically keeping the stability of the implicit solver, which will be explained in the paper. Another solver for the diffusion equation was found during the implementation of this project, namely a modified implicit solver. Ultimately, Dr. Bell and I settled on using the modified implicit solver as it has the same accuracy as the Crank-Nicolson solver while not being limited by the initial conditions being passed in.

Introduction

Atmospheres are extremely dynamic systems that require a lot of computations to accurately model. There are a lot of different interactions between molecules with different properties and densities. One major facet of atmospheric modeling is keeping track of the movement of energy. As a simplification, there are three big drivers in the movement of energy in an atmospheric system. These three forms of energy movement are convection, conduction, and radiation, namely thermal radiation. Convection is the movement of different compositions of molecules in an atmosphere. It is extremely dynamic which makes it very computationally expensive to model. Conduction is the collisions of molecules in an atmosphere, which can be extremely computationally expensive if one were to calculate the transfer of energy between individual molecules. However, this is not needed since conduction is can be macroscopically modelled with the diffusion equation. With such a powerful equation, there are a lot of ways to implement it to varying levels of accuracy. Unfortunately, this accuracy comes at a price of extra computational expenses. Lastly, thermal radiation is the emission of photons through the process of black body radiation. All things emit electromagnetic radiation based on an objects temperature, and planets are no different. Although convection and radiation are very important processes, this project's main focus is on the different implementations of calculating conduction through different methods of implementing the diffusion equation and how these different methods compare with each other.

To really highlight the differences of the these conduction models, every other atmospheric condition must be implemented. Given how difficult this would be to do, I will be using the Global Ionosphere Thermosphere Model, hereby referred to as GITM. GITM is an atmospheric framework used by the National Aeronautics and Space Administration (NASA). By using GITM, I am able to change only the conduction code and use the resultant data to numeri-

cally show the differences between the current conduction scheme used by GITM and two other implementations, the Crank Nicolson scheme and the Oscillation Free Second Order Implicit scheme.

Theory

Diffusion

As stated in the introduction, this line of research mainly focuses on the diffusion equation, which is given as

$$\frac{\partial U(\vec{r},t)}{\partial t} = \alpha \nabla(\kappa(\vec{r}) \nabla U(\vec{r},t)) \tag{1}$$

- *U* is the function of temperature with respect to time and position.
- ullet α contains all the constant factors such as constants intrinsic to the composition of the atmosphere
- \bullet κ contains constants of the atmosphere that are dependent on the distance from the planet such as density

This equation is currently written in three dimensional coordinates, but for this project, we will be exploring the diffusion equation is one dimension. This is achieved because the GITM framework uses parallel processes to run longitude and latitude blocks across multiple nodes in a cluster of computers. Therefore, conduction, which is modeled by the diffusion equation, is only run in one dimension on each of these blocks. As such, the diffusion equation is written as

$$\frac{\partial U(r,t)}{\partial t} = \alpha \frac{1}{r^2} \frac{\partial}{\partial r} (D(r) \frac{\partial U(r,t)}{\partial r})$$
 (2)

- As a simplification, D(r) is such that $D(r) = \kappa(r) * r^2$

This equations can further be simplified into two parts, a pseudo-Cartesian diffusion equation and a spherical correction, which are shown as:

$$\frac{\partial U(r,t)}{\partial t} = \alpha \frac{1}{r^2} \left(D(r) \frac{\partial^2 U(r)}{\partial r^2} + \frac{\partial D(r)}{\partial r} \frac{\partial U(r,t)}{\partial r} \right) \tag{3}$$

The following sections focus on the pseudo-Cartesian diffusion equation, and then the spherical correction will be explored. The reason this is a pseudo-Cartesian method comes from the inclusion of $\frac{1}{r^2}$ with the Cartesian diffusion equation.

Finite Difference of Diffusion

Although the diffusion equation is fairly simply solved, there is not an analytical solution for an atmosphere. Because of this, a new method is needed to implement this differential equation. This new method is finite difference. Finite difference is simply the decomposition of a differential equation into the difference between values at certain points over the difference between those points, simply written as:

$$\frac{dU(t)}{dt} = \frac{U^{n+1} - U^n}{\Delta t} \tag{4}$$

The notation for this equation is given as:

- \bullet *n* this is the step in time
- U^n corresponds to U(t)
- U^{n+1} corresponds to $U(t + \Delta t)$

This decomposition allows for using initial data to calculate future data, depending on the size of Δt . However, the greater the size of Δt , the less accurate the predicted value will be. As accuracy increases by reducing Δt , this increases the amount of iterations a processor must run the computation over again in order to reach a given time frame. As such, a balance between accuracy and computation time must be found.

Fortunately, reducing Δt is not the only way to improve accuracy. With every finite difference approximation, there is an associated order of accuracy. This order of accuracy is similar to the order of accuracy of a Taylor Series, as both are the error associated round off error. For the finite difference approximation given above, the associated error is first order accurate, which is written as $O(\Delta r)$. If a new method with higher accuracy is used, that computation time will only be increased marginally, by the approximation should have noticeable improvements to accuracy.

So far, the only finite difference equation explored has been for first order differential equations. While this is useful for decomposing the diffusion equation, a finite difference approximation is needed for second order differential equations. This decomposition is given as:

$$\frac{\partial^2 U(r)}{\partial r^2} = \frac{\frac{U_{j+1} - U_j}{\Delta r} - \frac{U_j - U_{j-1}}{\Delta r}}{\Delta r} \tag{5}$$

Simply put, a second order differential equation is the finite difference of a first order differential equation. Although this equation looks fairly messy, it can easily be rewritten as:

$$\frac{\partial^2 U(r)}{\partial r^2} = \frac{U_{j+1} - 2U_j + U_{j-1}}{\Delta r^2}$$
 (6)

The notation for this equation is given as:

- \bullet j this is the step in space
- U^j corresponds to U(r)
- U^{j-1} corresponds to $U(r \Delta r)$
- U^{j+1} corresponds to $U(r + \Delta r)$

One thing to note about this decomposition is that it is over Δr^2 . This is significant because this finite difference approximation is second order accurate.

With finite difference approximations for first order and second order differential equations, a finite difference approximation for the pseudo-Cartesian diffusion equation can be created. One such approximation would be:

$$\frac{U_j^{n+1} - U_j^n}{\Delta t} = \alpha \frac{1}{r^2} D_j \left(\frac{U_{j+1}^n - 2U_j^n + U_{j-1}^n}{\Delta r^2} \right)$$
 (7)

Some things to note with this approximation are

- U_i^{n+1} is what is trying to be solved
- $\bullet\,$ temporal differential equation is centered on U_j
- spatial differential equation is centered on U^n

This method of approximating the diffusion equation is called an Explicit scheme. This is because it explicitly uses information of the current time step to calculate the next time step. As such, it is fairly accurate. However, one major short coming of this scheme is that if $\Delta t \gg \Delta r^2/\alpha D_j$, the scheme will lose stability. As such, it requires a lot of computation time since Δt must be constrained to the constants and the size of the spatial mess. Therefore a new scheme is needed.

$$\frac{U_j^{n+1} - U_j^n}{\Delta t} = \alpha \frac{1}{r^2} D_j \left(\frac{U_{j+1}^{n+1} - 2U_j^{n+1} + U_{j-1}^{n+1}}{\Delta r^2} \right)$$
 (8)

This approximation is called the Implicit scheme, and it does not have the same short comings of the explicit scheme. As opposed to the implicit scheme, the spatial differential equation is centered on U^{n+1} . While this scheme is implying a break in casualty by using future data to calculate the current data, it is actually a way to set up the math of the system. Ultimately the system can be written as

$$\frac{U_j^{n+1}}{m'} - \frac{U_j^n}{m'} = D_j(U_{j+1}^{n+1} - 2U_j^{n+1} + U_{j-1}^{n+1})$$
(9)

Where $m' \equiv \frac{1}{r^2} \frac{\alpha \Delta t}{\Delta r^2}$. From here, all the \vec{U}^{n+1} are collected to one side and the \vec{U}^n are collected to the other side.

$$\frac{U_j^{n+1}}{m'} - D_j U_{j+1}^{n+1} + 2D_j U_j^{n+1} - D_j U_{j-1}^{n+1} = \frac{U_j^n}{m'}$$
 (10)

Therefore, there are j number of equations to form a $j \times j$ matrix of the form

$$\begin{bmatrix} \left(\frac{1}{m'} + 2D_{j}\right) & -D_{j} \\ -D_{j} & \left(\frac{1}{m'} + 2D_{j}\right) & -D_{j} \\ & \ddots & \ddots & \ddots \\ & & -D_{j} & \left(\frac{1}{m'} + 2D_{j}\right) & -D_{j} \\ & & & -D_{j} & \left(\frac{1}{m'} + 2D_{j}\right) \end{bmatrix}$$

$$(11)$$

If \widetilde{A} is this matrix, then this system of equations can be written as

$$\widetilde{A}\vec{U}^{n+1} = \vec{U}^n \tag{12}$$

With linear algebra, this matrix can be inverted to get a system of equations that can be implemented

$$\vec{U}^{n+1} = \tilde{A}^{n+1} \vec{U}^n \tag{13}$$

Non Uniform Grids

Before the spherical component can be analyzed, the subject of a non uniform grid must be addressed. The GITM framework uses a stretched mesh as altitude increases, so this must be accounted for. In all the talk above, I assumed a uniform mesh. Thankfully, the change to a non uniform mesh is rather simple for second order differential equations. Simply put, the finite difference approximation for second order differential equations becomes

$$\frac{\partial^2 U(r)}{\partial r^2} = \frac{\frac{U_{j+1} - U_j}{\Delta r_{j+1}} - \frac{U_j - U_{j-1}}{\Delta r_j}}{\frac{\Delta r_{j+1} + \Delta r_j}{2}} \tag{14}$$

- Δr_{j+1} is the value of r at U_{j+1} minus the value of r at U_j . $\Delta r_{j+1} = r_{j+1} r_j$
- Δr_j is the value of r at U_j minus the value of r at U_{j-1} . $\Delta r_j = r_j r_{j-1}$

This equation can be simplified to

$$\frac{\partial^2 U(r)}{\partial r^2} = \frac{1}{\frac{1}{2}(\Delta r_{j+1} + \Delta r_j)\Delta r_{j+1}\Delta r_j} (\Delta r_j U_{j+1} - (\Delta r_j + \Delta r_{j+1})U_j + \Delta r_{j+1}U_{j-1})$$
(15)

If $r' \equiv \frac{\Delta r_{j+1}}{\Delta r_j}$ then the equation can be simplified even further to

$$\frac{\partial^2 U(r)}{\partial r^2} = \frac{1}{\frac{1}{2}(\Delta r_{j+1} + \Delta r_j)\Delta r_{j+1}} (U_{j+1} - (1+r')U_j + r'U_{j-1})$$
(16)

With this, the implicit method described above becomes

$$\frac{U_j^{n+1}}{m} - \frac{U_j^n}{m} = \frac{D_j}{\frac{1}{2}(\Delta r_{j+1} + \Delta r_j)\Delta r_{j+1}} (U_{j+1}^{n+1} - (1+r')U_j^{n+1} + r'U_{j-1}^{n+1})$$
(17)

This equation still retains the second order accuracy explained in the above subsection, but now $\Delta r_{j+1} \neq \Delta r_j$. Further more, m' is no longer used, and instead m is used, which is given as $m \equiv \frac{\alpha \Delta t}{r^2}$.

Finite Difference of Spherical Correction

After analyzing the finite difference implementation of the pseudo-Cartesian diffusion equation, next is the spherical correction terms. Therefore, this section focuses on:

$$\frac{\partial D(r)}{\partial r} \frac{\partial U(r,t)}{\partial r} \tag{18}$$

Because these are all first order differential equations, this spherical correction could be approximated as

$$\frac{D_{j+1} - D_j}{\Delta r} \frac{U_{j+1}^{n+1} - U_j^{n+1}}{\Delta r} \tag{19}$$

One major fault of this approximation is that it is first order accurate. This is not ideal since the second order differential equation from the diffusion equation is second order accurate. It would be best to retain this accuracy if possible. Thankfully, there is a method to do this for first order differential equations, with a non uniform mesh. This method is given as:

$$\frac{dU(r)}{dr} = \frac{\Delta r_j}{\Delta r_{j+1}(\Delta r_j + \Delta r_{j+1})} U_{j+1} + \frac{\Delta r_{j+1} - \Delta r_j}{\Delta r_{j+1}\Delta r_j} U_j - \frac{\Delta r_{j+1}}{\Delta r_j(\Delta r_j + \Delta r_{j+1})} U_{j-1}$$
(20)

Using the notation that $r' \equiv \frac{\Delta r_{j+1}}{\Delta r_j}$, this equation can be simplified to

$$\frac{dU(r)}{dr} = \frac{1}{\Delta r_j r'(r'+1)} (U_{j+1} - (1 - r'^2)U_j - r'^2 U_{j-1})$$
(21)

With this approximation of a first order differential equation, the spherical correction can be approximated as

$$\frac{\partial D(r)}{\partial r} \frac{\partial U(r,t)}{\partial r} = \frac{1}{\Delta r_j r'(r'+1)}^2 (D_{j+1} - (1-r'^2)D_j - r'^2 D_{j-1})(U_{j+1} - (1-r'^2)U_j - r'^2 U_{j-1})$$
(22)

This equation can be simplified a little further if $dl_j \equiv D_{j+1} - (1 - r'^2)D_j - r'^2D_{j-1}$. This can be done because $\frac{\partial D(r)}{\partial r}$ is constant in time, so all that is needed is a spatial coordinate, denoted by the subscript j.

With the derivations of the pseudo-Cartesian and spherical correction equations, they can be combined to form the full implicit method that is used for the rest of this project.

$$\frac{U_j^{n+1}}{m} - \frac{U_j^n}{m} = \frac{D_j}{\frac{1}{2}(\Delta r_{j+1} + \Delta r_j)\Delta r_{j+1}} (U_{j+1}^{n+1} - (1+r')U_j^{n+1} + r'U_{j-1}^{n+1})$$
(23)

$$+\frac{dl_{j}}{(\Delta r_{j}r'(r'+1))^{2}}(U_{j+1}^{n+1}-(1-r'^{2})U_{j}^{n+1}-r'^{2}U_{j-1}^{n+1})$$
(24)

Methods

With the derivation of different finite difference schemes and a complete implementation of the implicit method of the diffusion equation, I want to shift into how this derivation will be applied to the GITM framework and the use of alternative methods to calculate diffusion. Since conduction is such a big part of energy and particle movement in atmospheres, it is extremely important to highlight the strengths and weaknesses of the different scheme and talk about their uses in GITM.

Implicit Method in Use

The current method used in the GITM framework is the implicit method explained above. As stated before, this method has second order accuracy in space, but only first order accuracy in time, which is not ideal. Below, I will explain two possible replacement schemes.

Crank-Nicolson Method

The Crank Nicolson scheme is an even combination of the Implicit and Explicit methods explained in the theory section. The reason I wanted to use the Crank-Nicolson scheme was that it has the unconditional stability of the implicit method while have more of the accuracy of the implicit method. Further more, this scheme maintains the second order accuracy of the implicit scheme while having second order accuracy in time because operations are performed on both the n and n+1 time steps. Naturally, this would seem to be an improved description of diffusion, however, there is an issue with this scheme when the initial values are very noisy, and it has to do with the amplification factor for this scheme, which is given as

$$G = \frac{1 - q(1 - \cos(\beta))}{1 + q(1 - \cos(\beta))}$$
 (25)

- $q = \frac{\alpha \Delta t}{\Delta r^2}$
- β is the Fourier frequency multiplied by Δr

This amplification factor is simply ratio of the amplitude of U^{n+1} compared to U^n .

$$\frac{U^{n+1}}{U^n} = \frac{1 - q(1 - \cos(\beta))}{1 + q(1 - \cos(\beta))}$$
 (26)

From this, the highest value of G can be 1, which means the system will have reached a steady state, since $U^n = U^{n+1}$. However, issues arise when G = -1. This means that the Crank-Nicolson scheme can allow oscillations to exist, which is problematic when $q \gg 1$. The main

tests Dr. Bell and I did in this experiment were on temperature via conduction, which does not exhibit the issues described here. However, diffusion is used in many other processes, one of which is viscosity. Unfortunately, he values of viscosity range so vastly that values where $q\gg 1$ do occur, and the model exhibits oscillatory behavior. For this reason, a newer method was needed.

Second Order Implicit Method

This newer method came in the form of a paper published by the Department of Mathematics and Applied Science at George Fox University. In their paper, titled "Oscillation-free method for semilinear diffusion equations under noisy initial conditions" [1], they describe an implicit method that uses a half step and full step to calculate the value of what a full step should be. This is done with the equation

$$U_W^{n+1} = 2 * U_H^{n+\frac{1}{2}} - U_F^{n+1} (27)$$

- ullet U_W^{n+1} is the weighed time step being calculated
- $\bullet \ \ U_H^{n+\frac{1}{2}}$ is the solution to an implicit scheme with only a half time step
- $\bullet \ U_F^{n+1}$ is the solution to an implicit scheme with a full time step

Since this method was recently published, Dr. Bell and I decided that it would be pertinent to first verify that this new method had the second order accuracy in time that was desired of the new method. There for Dr. Bell and I did a grid convergence on this method plotted against the first order implicit method used in the original GITM code.

As can be seen in figure 1, the new scheme has second order accuracy. With this condition satisfied, Dr. Bell and I decided to move to data collection with this new method as well as the Crank-Nicolson scheme to see if there were any differences between the methods outside of the oscillatory nature of the Crank-Nicolson.

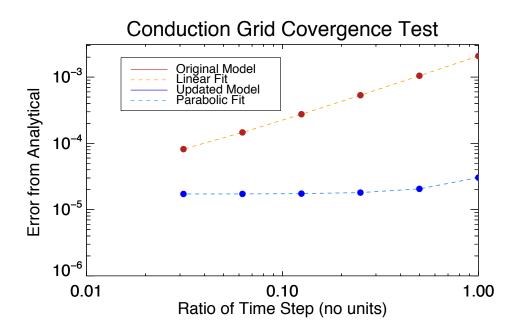


Figure 1: The equation being modeled is $E=Ch^p$, where E is the error, C is a constant, h is the step size, and p is the order of accuracy

Data From Mars

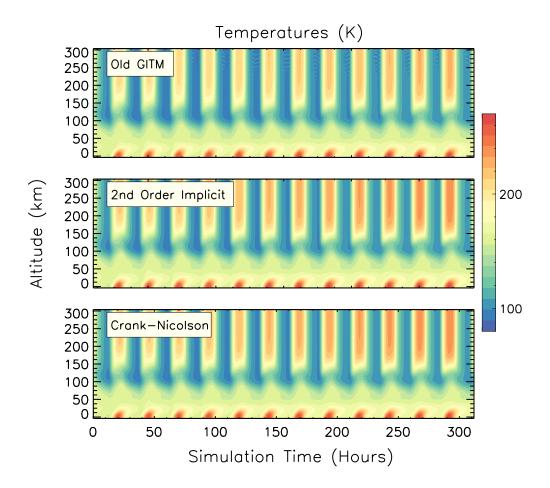


Figure 2: This is a stacked plot of the three simulations I ran. The first run, Old GITM, uses the Implicit scheme. The second run uses the Second Order Implicit scheme. The third run uses the Crank-Nicolson scheme.

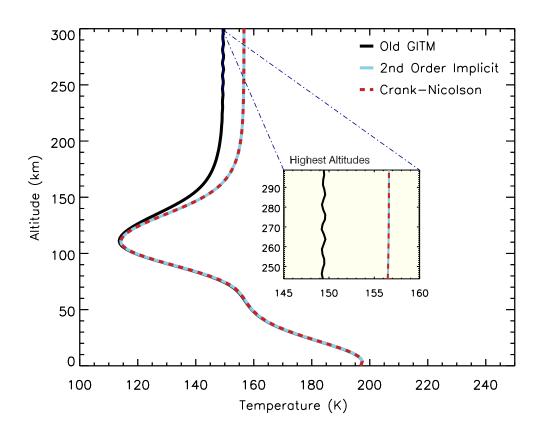


Figure 3: This is a snap shot from the stacked plots above, all compressed onto one grid for easy of viewing

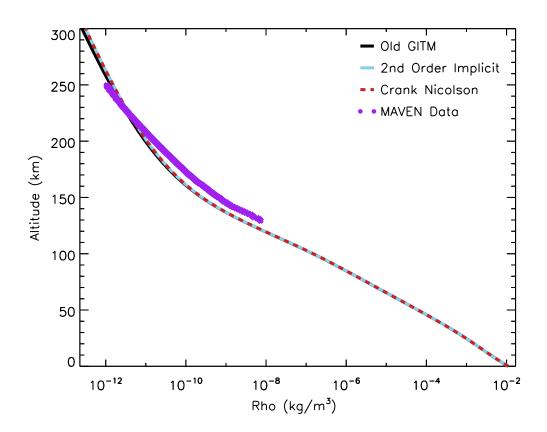


Figure 4: Here is a comparison of densities of the different runs compared to actual data taken from MAVEN, the mars orbiter

Discussion and Conclusions

From figure 2, it can be seen that there is a fairly significant difference between the original GITM code and the newer implementations. First, the temperature is higher in the upper atmosphere, as can be shown better in figure 3. Second, in figure 2, there is oscillation at the top of the Old GITM plot that seems to be propagating downward through the lifetime of the simulation. This was thought to be a quirk of the model itself, but it is actually the result of an improper implementation of a boundary condition at the top of the model. This boundary condition would switch the signs of the conduction part of the code as it was being calculated, such that $U^{n+1} = -U^n$ for the top cell of the atmosphere. The removal of this oscillatory nature was not due to the implementation of a newer scheme, but simply the implementation of a correct boundary condition.

In figure 3, it can be seen that there is a slight shift of the martian temperature profile at the top half of the atmosphere. This slight shift can be attributed to the higher accuracy of the scheme used. The oscillations of the old model have also been highlighted as they were another large change made to the model. One might notice that there is not model using the old boundary condition with a newer scheme. The reason for this is simply because the old boundary condition only created oscillations, but other than that, it did not have an effect on the temperature profile, so as to eliminate confusion, it was decided that the updated boundary condition would be used with the updated schemes.

In the last graph, there is a comparison to actual data taken from MAVEN, the martian orbiter created by NASA. As can be seen, the data better fits the data taken by MAVEN, however, the change is only 5%. This small of a change is nearly insignificant when looking at how far off from the actual data the models are. Although, it does not align perfectly with the data, I want to point out that the data being looked at is density, and not temperature. For this reason,

this is not quite an accurate metric of the accuracy of the new model. Further more, the data used is taken from the planet, which is a 3 dimensional entity, while these simulations are 1 dimensional. Although running a 3 dimensional model will not effect the temperature, it will effect the density distribution since there will be convection blowing around the gases in Mars' atmosphere.

Overall, the goal of this project was to increase the accuracy conduction code used in the GITM framework and to show that there was a measurable difference. By using the Crank-Nicolson and Second Order Implicit schemes, I showed that the difference between these schemes and the original implicit scheme is due to the increased accuracy, since these schemes would not yield the same results when used to simulate the atmosphere of Mars if it wasn't due to the increased order of accuracy.

Appendices

In this section, I present the Earth data I took earlier in the life of the project. This data was left unused in favor of the Mars data due to the lack of evidence of any significant change between the Crank-Nicolson scheme and the Implicit scheme. This failing is not attributed to the Crank-Nicolson scheme, but to the atmospheric dynamics of Earth. Conduction is the main atmospheric process that uses the diffusion equations, so if I wanted to highlight the differences between schemes that implement the diffusion equation, I would want an atmosphere that has a large part of its atmospheric dynamics coming from conduction. For this reason, Mars was the prime candidate of study for this project. With this in mind, there is change between the Crank-Nicolson scheme and the Implicit scheme at the boundary conditions and with some of the highs and lows of the atmosphere, but Dr Bell and I felt that we could get better measurements with Mars.

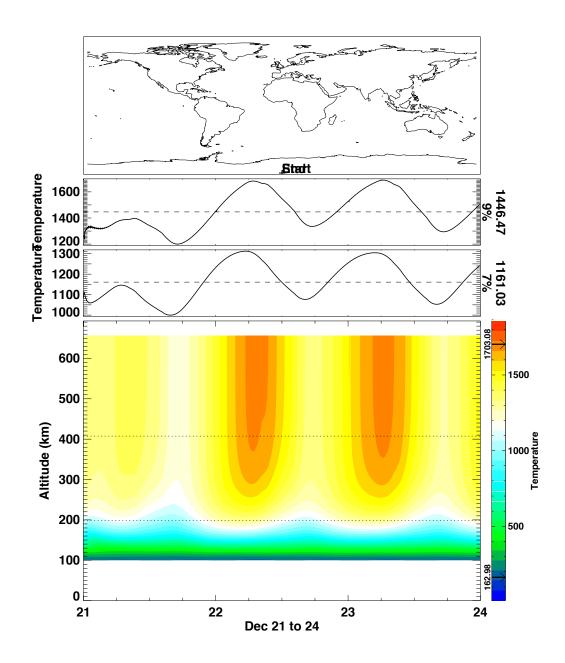


Figure 5: This is data from an Earth simulation using the original GITM code

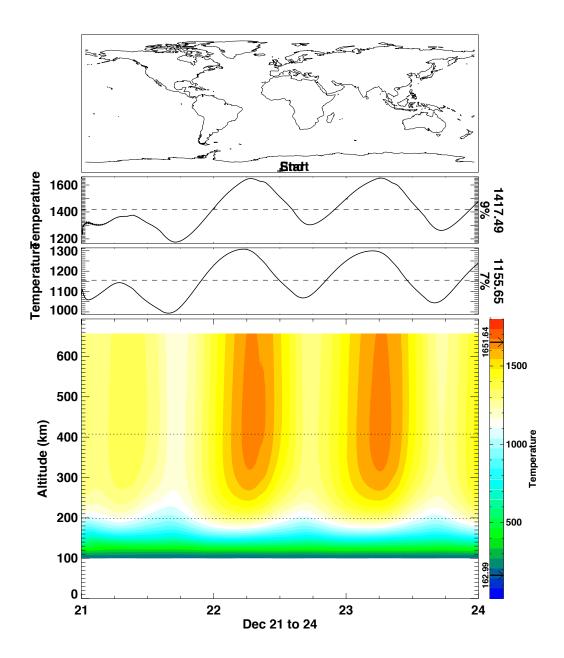


Figure 6: This is data from an Earth simulation using the Second Order Implicit scheme

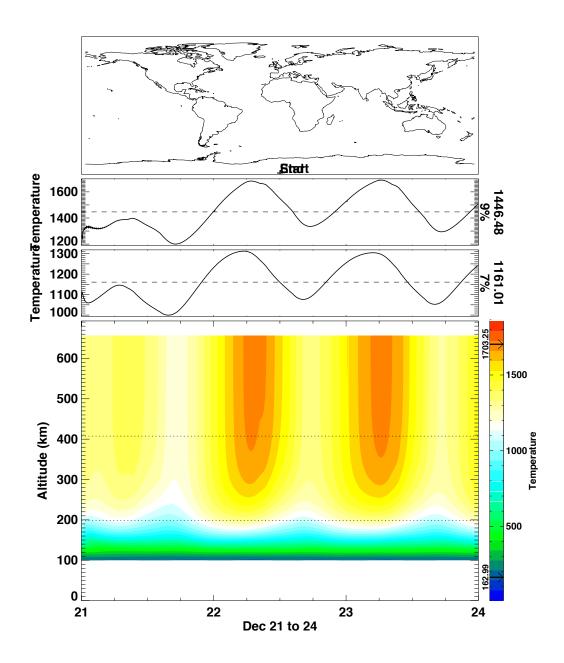


Figure 7: This is data from an Earth simulation using the Crank-Nicolson scheme

Bibliography

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initial conditions" (2016). Faculty Publications - Department of Mathematics and Applied Science. Paper 16.

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